

Height fluctuations for the stationary KPZ equation

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In their seminal 1986 paper [13], Kardar, Parisi, and Zhang (KPZ) proposed the stochastic evolution equation for a height function $h(t, x) \in \mathbb{R}$ ($t \in \mathbb{R}_+$ is time and $x \in \mathbb{R}$ is space)

$$\partial_t h(t, x) = \frac{1}{2} \partial_x^2 h(t, x) + \frac{1}{2} (\partial_x h(t, x))^2 + \xi(t, x).$$

The randomness ξ models the deposition mechanism and it is taken to be space-time Gaussian white noise, so that formally $\mathbb{E}[\xi(t, x)\xi(t', x')] = \delta(t - t')\delta(x - x')$. The Laplacian reflects the smoothing mechanism and the non-linearity reflects the slope-dependent growth velocity of the interface. Using earlier physical work of Forster, Nelson and Stephen [10] KPZ predicted that for large time t , the height function $h(t, x)$ has fluctuations around its mean of order $t^{1/3}$ with spatial correlation length of order $t^{2/3}$. For additional background, see the reviews [8, 9, 15, 17].

The physically relevant solution to the KPZ equation [17] is defined indirectly via the well-posed stochastic heat equation (SHE) with multiplicative noise [2, 5, 16],

$$\partial_t \mathcal{Z}(t, x) = \frac{1}{2} \partial_x^2 \mathcal{Z}(t, x) + \mathcal{Z}(t, x) \xi(t, x),$$

with initial condition $\mathcal{Z}(0, x) = \mathcal{Z}_0(x) = e^{h(0, x)}$. The SHE is well-posed and we defines $h(t, x) = \ln(\mathcal{Z}(t, x))$. This is called the Cole–Hopf solution of the KPZ equation.

By a version of the Feynman–Kac formula, the solution of the SHE can be written as

$$\mathcal{Z}(t, x) = \mathbb{E}_{t, x} \left[\mathcal{Z}_0(b(0)) : \exp : \left(- \int_0^t \xi(b(s), s) ds \right) \right]$$

where the expectation $\mathbb{E}_{t, x}$ is over a Brownian motion $b(\cdot)$ going backwards in time from $b(t) = x$, and where $: \exp :$ is the Wick ordered exponential. This provides an interpretation for $\mathcal{Z}(t, x)$ as the partition function of the continuum directed random polymer (CDRP) [1, 2].

Let $B(x)$ be a two-sided Brownian motion with $B(0) = 0$ and zero drift. Stationary (zero drift) initial data $h(0, x) = B(x)$ for the KPZ equation corresponds with SHE initial data $\mathcal{Z}(0, x) = e^{B(x)}$. This is called stationary because for any later time t , $h(t, \cdot)$ is marginally distributed as $\tilde{B}(\cdot) + h(t, 0)$ where $\tilde{B}(\cdot)$ is a two-sided Brownian motion (though not independent of B or $h(t, 0)$).

In our work [7] we provide an exact formula for the one-point probability distribution of the stationary solution to the KPZ equation, and a limit theorem for $h(t, x)$, after proper centering and scaling by $t^{1/3}$. This is made by analyzing a semidiscrete directed polymer model, which in its turn is obtained as a limit of the q-Whittaker process [6] with appropriate initial measure. A different expression

using replica trick approach was obtained previously in [12]. The equivalence of the two formulas has not been shown so-far.

For simplicity, let us present the results in the case of zero-drift and position $x = 0$ only.

Theorem 1. *Let $h(t, x)$ be the stationary (zero drift) solution to the KPZ equation and let K_0 denote the modified Bessel function. Then, for $t > 0$, $\sigma = (2/t)^{1/3}$ and $S \in \mathbb{C}$ with positive real part,*

$$\mathbb{E} \left[2\sigma K_0 \left(2\sqrt{S \exp \left[\frac{t}{24} + h(t, 0) \right]} \right) \right] = f(S, \sigma),$$

where the function f is given below.

To define f , define on \mathbb{R}_+ the function

$$Q(x) = \frac{1}{2\pi i} \int_{-\frac{1}{4\sigma} + i\mathbb{R}} dw \frac{\sigma \pi S^{-\sigma w}}{\sin(\pi - \sigma w)} e^{-w^3/3 + wx} \frac{\Gamma(\sigma w)}{\Gamma(-\sigma w)},$$

and the kernel

$$K(x, y) = \frac{1}{(2\pi i)^2} \int_{-\frac{1}{4\sigma} + i\mathbb{R}} dw \int_{\frac{1}{4\sigma} + i\mathbb{R}} dz \frac{\sigma \pi S^{\sigma(z-w)}}{\sin(\sigma \pi(z-w))} \frac{e^{z^3/3 - zy}}{e^{w^3/3 - wx}} \frac{\Gamma(-\sigma z)}{\Gamma(\sigma z)} \frac{\Gamma(\sigma w)}{\Gamma(-\sigma w)}.$$

Let $\gamma_E = 0.577\dots$ be the Euler constant, $\gamma = \sigma(2\gamma_E + \ln S)$ and define

$$f(S, \sigma) = -\det(\mathbb{1} - K) \left[\gamma + \langle (\mathbb{1} - K)^{-1}(K\mathbb{1} + Q), \mathbb{1} \rangle + \langle (\mathbb{1} - K)^{-1}(\mathbb{1} + Q), Q \rangle \right].$$

where the determinants and scalar products are all meant in $L^2(\mathbb{R}_+)$.

There is an explicit inversion formula, although this is not needed to get the large time limit.

Corollary 2. *For any $r \in \mathbb{R}$, we have*

$$\begin{aligned} & \mathbb{P} \left(h(t, 0) \leq -\frac{t}{24} + r \left(\frac{t}{2} \right)^{1/3} \right) \\ &= \frac{1}{\sigma^2} \frac{1}{2\pi i} \int_{-\delta + i\mathbb{R}} \frac{d\xi}{\Gamma(-\xi)\Gamma(-\xi+1)} \int_{\mathbb{R}} dx e^{x\xi/\sigma} f \left(e^{-\frac{x+r}{\sigma}}, \sigma \right) \end{aligned}$$

for any $\delta > 0$ and where $\sigma = (2/t)^{1/3}$.

Universality arises in the large time limit. Indeed, we recover the distribution formula obtained before for the totally asymmetric simple exclusion process and for the polynuclear growth model [3, 4, 11, 14].

Corollary 3. *For any $r \in \mathbb{R}$,*

$$\lim_{t \rightarrow \infty} \mathbb{P} \left(h(t, 0) \leq -\frac{t}{24} + r \left(\frac{t}{2} \right)^{1/3} \right) = F_0(r),$$

where F_0 is given by

$$F_0(r) = \frac{\partial}{\partial r} \left(g(r) \det(\mathbb{1} - P_r K_{A_i} P_r)_{L^2(\mathbb{R})} \right),$$

where $P_s(x) = \mathbb{1}_{\{x>s\}}$, $K_{\text{Ai}}(x, y) = \int_0^\infty d\lambda \text{Ai}(x + \lambda)\text{Ai}(y + \lambda)$ is the Airy kernel, and $g(r)$ is given below.

To define the function g we need some notations. For $s \in \mathbb{R}$, define

$$\begin{aligned} \mathcal{R} &= s + \int_s^\infty dx \int_0^\infty dy \text{Ai}(x + y), & \Psi(y) &= 1 - \int_0^\infty dx \text{Ai}(x + y), \\ \Phi(x) &= \int_0^\infty d\lambda \int_s^\infty dy \text{Ai}(x + \lambda)\text{Ai}(y + \lambda) - \int_0^\infty dy \text{Ai}(y + x). \end{aligned}$$

Then the function g is defined by

$$g(s) = \mathcal{R} - \langle (\mathbb{1} - P_s K_{\text{Ai}} P_s)^{-1} P_s \Phi, P_s \Psi \rangle.$$

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